

STATE AGENCY ON SCIENCE, INNOVATION
AND INFORMATION OF UKRAINE
RESEARCH INSTITUTE “NIKOLAEV ASTRONOMICAL OBSERVATORY”

**ASTRONOMICAL RESEARCH:
FROM NEAR-EARTH SPACE
TO THE GALAXY**

International Conference

ABSTRACT BOOK

September 26-29, 2011,
Mykolaiv, Ukraine

d) calculations of the orbit and brightness for small objects using celestial mechanics algorithms;

- powerful data search engine;
- customizable format and logic control for linkage of the local virtual observatory with other similar data storages, as well as ingressed into IVOA Alliance, or independent astronomical data archives of selected observatories.
- 2D and 3D visualization of the objects.

Usage of .NET technologies is in the core of this software complex. Complex supports Oracle, Microsoft SQL Server, Sybase, PostgreSQL, FireBird, MySQL DBMSs. As an HTTP server for this complex could be used such servers like Apache, or Microsoft IIS. Complex "Adelaide" runs under OS Windows, Linux or Android. Also, software "Adelaide" can transform data storage for using it in cloud services.

NIKOLAEV OBSERVATORY IN THE FIRST HALF OF XX CENTURY

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At the beginning of last century the tremendous changes expected Nikolaev Observatory, which was subordinated to Navy Ministry during 90 years. The main event of this period was its transformation in 1912 into one of the southern department of Pulkovo Observatory with the aim to create a base for astrometric researches in the South. There was the time of the greatest upheavals in our society: the First World War, the Revolution of 1917, the Civil War. A small respite, which was associated with the revival of the observatory in 20-30 years of XX century, was interrupted by new troubles - the Great Patriotic War. The Observatory have survived all these troubles, and have preserved as a scientific institution, and as a historical and architectural complex. Of course, the credit for this was due to the two directors of the observatory - Boris Pavlovich Ostaschenko-Kudryavtsev (1909-1923) and Leonid Ivanovich Semenov (1923-1950), who came from the Pulkovo Observatory. During those years, the observatory participated in the first major international project on developing of the sky maps. Absolute catalogs of stars, which were

compiled from observations with Freiberg-Kondratiev transit instrument and Repsold vertical circle, were included as a part in the fundamental catalogs FK. In 1931 the Time service, which was one of the best in the USSR, was organized in the Observatory. Another target of Nikolaev department was to determine the positions of the Sun and Solar system bodies.

Currently, we are working on a new book of the series of biobibliographical books devoted to the directors of NAO - B.P. Ostaschenko-Kudryavtsev and L.I. Semenov. Materials in the archives of the NAO, Pulkovo Observatory and personal archives of B.P. Ostaschenko-Kudryavtseva are used to write it.

THE INFLUENCE OF YARKOVSKY AND YORP EFFECTS ON DYNAMICAL EVOLUTION OF ASTEROIDS

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The two effects are produced by a weak non-gravitational acceleration of asteroids and meteoroids, which is caused by radiative recoil due to anisotropic thermal re-emission of energy. The first one produces secular changes in orbital semi-major axis of body and is called Yarkovsky effect. The second one produces changes in rotational state and is called YORP (Yarkovsky-O'Keefe-Radzievskiy-Paddack) effect. It is believed that these effects have significant influence on orbital and rotational dynamics of asteroids less than 100 km. The Yarkovsky and YORP effects were not investigated earlier due to insufficient precision of observations.

The Yarkovsky effect was predicted in 1900 and firstly confirmed in 2003 up to now only for one asteroid (6489) Golevka using radar ranging. At present, YORP effect was detected for 4 asteroids using photometric data. All of these asteroids are NEAs, and detection of YORP and Yarkovsky effects for main belt asteroids is forthcoming.

The review presents basic principles of the Yarkovsky and YORP effects, their possible influences on dynamical evolution of asteroids and recent results. The estimation and inclusion these effects to a model of motion among with other already used factors will improve our understanding of the dynamical evolution of small Solar system bodies.