

ABSTRACT BOOK



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S236-101 Poster

Astrometric Observations Of Asteroids And Comets At The Moletai Astronomical Observatory

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We describe the observational project devoted to astrometric observations of Near-Earth Objects (NEO), main belt asteroids and comets at the Moletai Observatory, Lithuania. Exposures are obtained with the two telescopes of the observatory: 0.35/0.50-m f/3.5 Maksutov telescope and the 1.65-m f/3.1 reflector with a CCD camera. The results of more than 10000 positions of asteroids and comets have been published in the Minor Planet Circulars and Minor Planet Electronic Circulars. During the 2001–2006 period 125 new main belt asteroids were discovered and a few NEO objects were detected independently.

S236-102 Poster

Use of combined CCD method of observations for NEO in Nikolaev observatory

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The number of discovered by astronomers asteroids in tens and hundreds meters is increasing every year. Several hundreds of such bodies approach the Earth to 0.1 a.u. during the year. Use of small telescopes ($D \approx 1\text{m}$) will significantly increase statistics of observations and accordingly improve prediction of asteroid's orbit. Combined CCD method of observations, developed in NAO, was tested for observations of NEO. The method is realized with the use of special mechanical device, the rotary platform. The rotary platform provides realization of drift-scan mode along the direction of visible motion of NEO, and for reference stars in their diurnal motion. Observations were made at such telescopes, as "Fast robotic telescope" ($D = 300\text{mm}$, $F = 1500\text{ mm}$) of Nikolaev observatory and AZT-8 ($D = 700\text{ mm}$, $F = 2800\text{ mm}$) of National control center and test of space means in Evpatoria. There were observed such NEO, as 2004WG1, 2005TF49, 2001SG276. The use of combined method has allowed to increase in 2.5 limiting magnitudes of telescopes. Accuracy of a single observation was $0.1 - 0.3$. The accuracy of observations at the telescopes are given and the values of (O-C) from comparison with MPC data. It is possible to conclude, that the developed and realized in NAO combined method of observations significantly expands possibilities of small telescopes for observations of potentially hazardous asteroids for the Earth.

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Solar System Bodies "Observations in the Past" with Plate Archive of Main Astronomical Observatory of Ukrainian NAS.

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The improvement of the dynamical models of Solar System bodies' motions will be very useful for the future space astrometry mission GAIA for a fast identification of objects to discriminate between the well-known objects and the new ones. "Observations in the Past" with Plate Archives allow realising it.

The Plate Archive of the Main Astronomical Observatory of the National Academy of Sciences of Ukraine contains more than hundred thousand images of minor planets with magnitude up to $16^m.7$. About 10 % of minor planets, which may be found on our archive plates, were firstly discovered after time when plates have been taken. So we can rediscovery them by so called "observation in the past" and obtain their positions.

Now we intend to find, measure and define of orbital elements for: "rediscovered" asteroids, asteroids with out-of-truth orbits and Near Earth Objects. Other Solar System bodies for which we try to get those "observation in the past" are external planets' satellites.

The objects choose criteria, methods of it search, identification determination of position are discussed. The results of external planets' satellites search in MAO plate archive will be

S236-104 Poster

New Methods for the Study of the Motion of Asteroids and

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A review of the new effective methods based on a superoscillating intermediate orbits developed by the author. These methods are meant for the following three main problems of dynamical astronomy: 1) preliminary orbit determination; 2) initial parameters of motion; 3) high accuracy prediction of positions and velocities of celestial bodies.

For the solution of the first problem a new method of finding a celestial body with the use of two position vectors of this corresponding time interval is constructed. A method is proposed for determining a preliminary orbit from three positions of a celestial sphere that was developed by analogy with Lagrange–Gauss method.

To solve the second problem new algorithms are proposed for computation of partial derivatives of the current parameters respect to their initial values using the arcs of superoscillating intermediate orbits.

The applications of the orbits constructed in Encke's method perturbations and in the procedure for predicting the motion of perturbed trajectory is represented by a sequence of superoscillating intermediate orbits are suggested for solving the third problem. The first application leads to generalized Encke's method. The second — to new methods for solving the equations of motion whose accuracy order coincides with the tangency of the superoscillating intermediate orbit used. The application of the Runge rule and extrapolation allow us to obtain methods of higher orders.

The estimations of efficiency of the developed methods in the case of investigation of motion of some minor planets and comets. It is shown that the new methods raise the accuracy of the operation of machine computations as compared with analogous methods.

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Initial Linking Methods and their Classification.

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The problem of initial linking of asteroids is of increasing importance for the next generation surveys. During the first week after discovery the orbital elements are very uncertain and other methods are used. In this paper 8 initial linking methods are given. There are two different types of search area is computed on a second night from the first night. The undoubtedly linked positions, typically on the first night. This assumes candidates which are then checked by the comparison of residuals of an orbit. Computations may be classified as being in 3-dimensional space or the 2-dimensional sky-plane.

A new basis, with a simpler computational algorithm, is proposed. The widely used Väisälä method. For a new method a simple programme is given.